

Coherence and Phase-space II

VSSUP Lectures 2014

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January 21, 2014

Outline

- 1 Quantum dynamics
- 2 Exponential complexity
- 3 Wigner stochastic equations

Ultracold atoms - the ideal quantum system

ULTRALOW temperatures down to $1nK$

What is different about ultracold atoms?

- Atoms are trapped in a hard vacuum
- Cooling to nanoKelvins or less
- Can have either bosons or fermions
- Atom 'lasers' - atoms behave as quantum objects
- Correlations - mean field theory doesn't always work
- Dynamics - time-evolution is very important

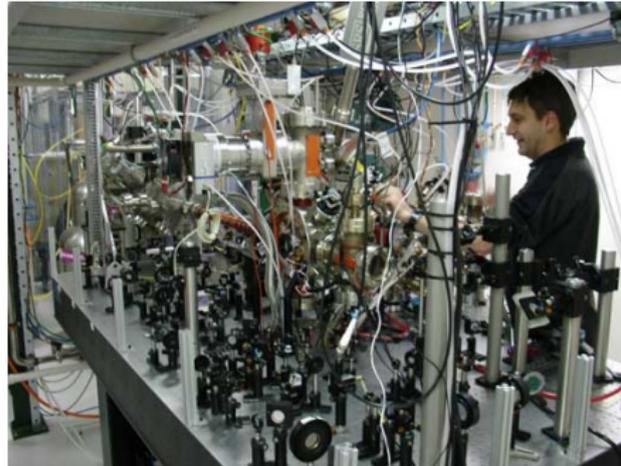
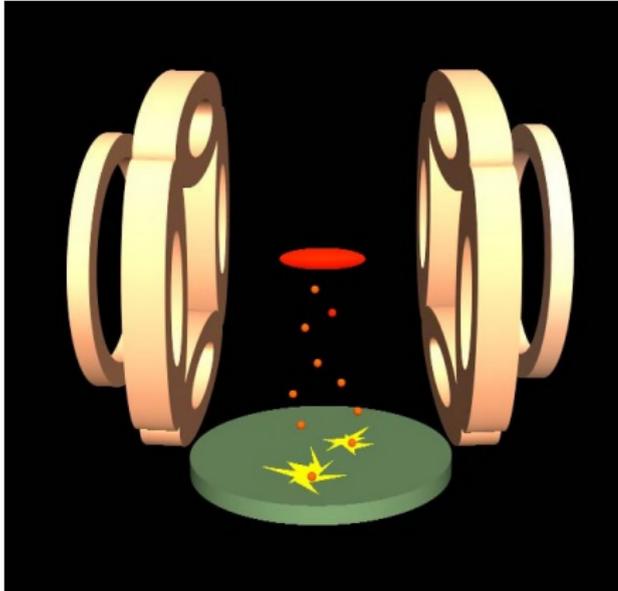
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Typical experiment (Orsay, ANU)



How to calculate dynamics?

Classical solution: - use Hamilton's equations

$$\dot{p}_i = -\frac{\partial H}{\partial q_i}$$

$$\dot{q}_i = \frac{\partial H}{\partial p_i}$$

Quantum mechanics replaces classical quantities by corresponding operators with commutators, so that

$$[\hat{q}_i, \hat{p}_j] = i\hbar\delta_{ij}$$

$$[\hat{q}_i, \hat{q}_j] = [\hat{p}_i, \hat{p}_j] = 0$$

Then, for any operator \hat{O} , in the Heisenberg picture:

$$\frac{\partial \hat{O}}{\partial t} = \frac{1}{i\hbar} [\hat{O}, \hat{H}]$$

What about mixtures of states?

Suppose the quantum system is in a mixture of quantum states $|\psi_m\rangle$ with probability p_m . Then the density matrix $\hat{\rho}$ is defined as:

$$\hat{\rho} = \sum_m p_m |\psi_m\rangle \langle \psi_m|$$

In the Schroedinger picture, we let states evolve in time, not operators!

$$\frac{\partial \hat{\rho}}{\partial t} = \frac{1}{i\hbar} [\hat{H}, \hat{\rho}]$$

Then, for any operator \hat{O} , the expectation value of the observable is:

$$\langle \hat{O} \rangle = \text{Tr} [\hat{\rho} \hat{O}]$$

What is the Hamiltonian anyway?

What about the quantum fields with hats?

Here, $\hat{\Psi}_i$ is a quantum field of spin-index i :

$$\left[\hat{\Psi}_i(\mathbf{x}), \hat{\Psi}_j^\dagger(\mathbf{x}') \right]_{\pm} = \delta_{ij} \delta^D(\mathbf{x} - \mathbf{x}')$$

In second quantization the quantum Hamiltonian is

$$\begin{aligned} \hat{H} &= \sum_i \int d^D \mathbf{x} \left\{ \frac{\hbar^2}{2m} \nabla \hat{\Psi}_i^\dagger(\mathbf{x}) \cdot \nabla \hat{\Psi}_i(\mathbf{x}) + V_i(\mathbf{x}) \hat{\Psi}_i^\dagger(\mathbf{x}) \hat{\Psi}_i(\mathbf{x}) \right\} \\ &+ \sum_{ij} \frac{U_{ij}}{2} \int d^D \mathbf{x} \hat{\Psi}_i^\dagger(\mathbf{x}) \hat{\Psi}_j^\dagger(\mathbf{x}) \hat{\Psi}_j(\mathbf{x}) \hat{\Psi}_i(\mathbf{x}) . \end{aligned}$$

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What are the parameters?

This describes a dilute gas at low enough temperatures,

- $\langle \hat{\Psi}_i^\dagger(\mathbf{x}) \hat{\Psi}_i(\mathbf{x}) \rangle$ is the spin i atomic density,
- m is the atomic mass,
- V_i is the atomic trapping potential & Zeeman shift
- U_{ij} is related to the S-wave scattering length in three dimensions by:

$$U_{ij} = \frac{4\pi\hbar^2 a_{ij}}{m}.$$

- Here we implicitly assume a momentum cutoff $k_c \ll 1/a$

Simplest method for state evolution

Suppose the quantum system is described by a few modes:

$$|\psi\rangle = \sum \psi_{\mathbf{N}} |N_1, N_2, \dots, N_m\rangle = \sum \psi_{\mathbf{N}} |\mathbf{N}\rangle$$

Then, let $H_{\mathbf{NM}} = \langle \mathbf{N} | \hat{H} | \mathbf{M} \rangle$ and: $\frac{d}{dt} |\psi\rangle = -\frac{i}{\hbar} \hat{H} |\psi\rangle$
 Hence, we have a simple matrix equation:

$$\frac{d}{dt} \psi_{\mathbf{N}} = -\frac{i}{\hbar} \sum_{\mathbf{M}} H_{\mathbf{NM}} \psi_{\mathbf{M}}$$

(4) Prove the last equation using orthogonality

Problem: quantum theory is exponentially complex!

Quantum many-body problems are very large

- consider N particles distributed among M modes
- take $N \simeq M \simeq 500,000$:
- Number of quantum states: $N_s = 2^{2N} = 2^{1,000,000}$
- More quantum states than atoms in the universe
- How big is your computer?
- **Can't diagonalize $2^{1,000,000} \times 2^{1,000,000}$ Hamiltonian!**

What about losses and damping?

Damping can be treated using a master equation

- The density matrix $\hat{\rho}$ evolves as:

$$\frac{d\hat{\rho}}{dt} = -\frac{i}{\hbar} [\hat{H}, \hat{\rho}] + \sum_j \kappa_j \int d^3\mathbf{r} \mathcal{L}_j[\hat{\rho}]$$

- Here the Liouville terms describe coupling to the reservoirs:

$$\mathcal{L}_j[\hat{\rho}] = 2\hat{O}_j\hat{\rho}\hat{O}_j^\dagger - \hat{O}_j^\dagger\hat{O}_j\hat{\rho} - \hat{\rho}\hat{O}_j^\dagger\hat{O}_j$$

- For n-particle collisions: $\hat{O}_i = [\hat{\Psi}_i(\mathbf{r})]^n$

Traditional quantum theory methods?

- numerical diagonalisation?
intractable for $\gtrsim 10$ modes
- operator factorization
not applicable for strong correlations
- perturbation theory
diverges at strong couplings
- exact solutions
not applicable for quantum dynamics

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Quantum theory in classical phase-space

Properties of Wigner/Moyal phase-space

- Maps quantum states into **classical phase-space** $\alpha = p + ix$
- **Wigner first published this representation**
- Moyal showed equivalence to quantum mechanics
- **Complexity grows only linearly with number of modes!**

Problem: Wigner distribution can have negative values

- **Need to truncate equations to get positive probabilities**

Detailed equivalence

Mapping of characteristic functions

$$W(\boldsymbol{\alpha}) = \frac{1}{\pi^{2M}} \int d^{2M} \mathbf{z} \left\langle e^{i\mathbf{z} \cdot (\hat{\mathbf{a}} - \boldsymbol{\alpha}) + iz^* \cdot (\hat{\mathbf{a}}^\dagger - \boldsymbol{\alpha}^*)} \right\rangle$$

Operator mean values

- $\langle \hat{a}_i^{\dagger m} \hat{a}_j^n \rangle_{SYM} = \int d^{2M} \boldsymbol{\alpha} \alpha_i^{*m} \alpha_j^n W(\boldsymbol{\alpha}) = \langle \alpha_i^{*m} \alpha_j^n \rangle_W$
- $\langle \hat{a}_j \rangle = \langle \alpha_j \rangle_W$
- $\langle \hat{a}_j^\dagger \hat{a}_j + \hat{a}_j \hat{a}_j^\dagger \rangle / 2 = \langle \alpha_j^* \alpha_j \rangle_W$

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Dynamical equivalence

Mapping of dynamical equations

$$\frac{\partial W(\boldsymbol{\alpha})}{\partial t} = \frac{1}{\pi^{2M}} \int d^{2M} \mathbf{z} \text{Tr} \left[\frac{\partial \hat{\rho}}{\partial t} e^{i\mathbf{z} \cdot (\hat{\mathbf{a}} - \boldsymbol{\alpha}) + i\mathbf{z}^* \cdot (\hat{\mathbf{a}}^\dagger - \boldsymbol{\alpha}^*)} \right]$$

Operator mappings

- $\hat{a}_j \hat{\rho} \rightarrow \left(\alpha_j + \frac{1}{2} \frac{\partial}{\partial \alpha_j^*} \right) W$
- $\hat{\rho} \hat{a}_j^\dagger \rightarrow \left(\alpha_j^* + \frac{1}{2} \frac{\partial}{\partial \alpha_j} \right) W$
- $\hat{a}_j^\dagger \hat{\rho} \rightarrow \left(\alpha_j^* - \frac{1}{2} \frac{\partial}{\partial \alpha_j} \right) W$
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Example: Wigner function for a coherent state

Suppose we have a single-mode BEC in a coherent state

$$\hat{\rho} = |\alpha_0\rangle\langle\alpha_0|$$

Hence:

$$W(\alpha) = \frac{1}{\pi^2} \int d^2z \langle\alpha_0| e^{iz\cdot(\hat{a}-\alpha)+iz\cdot(\hat{a}^\dagger-\alpha^*)} |\alpha_0\rangle$$

Solution with a little algebra

$$W(\alpha) = \frac{2}{\pi} e^{-2|\alpha-\alpha_0|^2}$$

(5): show that this solution gives $\langle\alpha^*\alpha\rangle = 1/2$ for a vacuum state

Example: time-evolution of harmonic oscillator

Consider the harmonic oscillator

$$\hat{H} = \hbar\omega \hat{a}^\dagger \hat{a}$$

$$\frac{\partial \hat{\rho}}{\partial t} = -i\omega [\hat{a}^\dagger \hat{a} \hat{\rho} - \hat{\rho} \hat{a}^\dagger \hat{a}]$$

Operator mappings

- $\hat{a}^\dagger \hat{a} \hat{\rho} \rightarrow \left(\alpha^* - \frac{1}{2} \frac{\partial}{\partial \alpha} \right) \left(\alpha + \frac{1}{2} \frac{\partial}{\partial \alpha^*} \right) W$

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$$\frac{\partial W}{\partial t} = i\omega \left(\frac{\partial}{\partial \alpha} \alpha - \frac{\partial}{\partial \alpha^*} \alpha^* \right) W$$

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Harmonic oscillator solution

General result for harmonic oscillator

$$\frac{\partial W}{\partial t} = i\omega \left(\frac{\partial}{\partial \alpha} \alpha - \frac{\partial}{\partial \alpha^*} \alpha^* \right) W$$

Solution by method of characteristics



$$\frac{\partial \alpha}{\partial t} = -i\omega \alpha$$



$$\alpha(t) = \alpha(0)e^{-i\omega t}$$

(6): Prove this!

Fokker-Planck equations

Result of operator mappings:

$$\frac{\partial W}{\partial t} = \left\{ -\frac{\partial}{\partial \alpha_i} A_i + \frac{1}{2} \frac{\partial^2}{\partial \alpha_i \partial \alpha_j^*} D_{ij} + \frac{1}{6} \frac{\partial^3}{\partial \alpha_i \partial \alpha_j^* \partial \alpha_k^*} T_{ijk} + \dots \right\} W$$

Scaling to eliminate higher-order terms

$$x = \alpha / \sqrt{N}$$

$$\frac{\partial W}{\partial t} = \left\{ -\frac{1}{\sqrt{N}} \frac{\partial}{\partial x_i} A_i + \frac{1}{2N} \frac{\partial^2}{\partial x_i \partial x_j} D_{ij} + O\left(\frac{1}{N^{3/2}}\right) \right\} W$$

Stochastic equation

Result of operator mappings + truncation - valid if $N/M \gg 1$:

$$\frac{\partial W}{\partial t} = \left\{ -\frac{\partial}{\partial \alpha_i} A_i + \frac{1}{2} \frac{\partial^2}{\partial \alpha_i \partial \alpha_j^*} D_{ij} \right\} W$$

Equivalent stochastic equation

$$\frac{\partial \alpha_i}{\partial t} = A_i + \zeta_i(t)$$

where:

$$\langle \zeta_i(t) \zeta_j^*(t') \rangle = D_{ij} \delta(t - t')$$

Example: BEC case

Result of operator mappings + truncation - for the GPE:

$$\frac{d\psi_j}{dt} = iK_j\psi_j - iU_{ij}|\psi_i|^2\psi_j - \gamma_j\psi_j + \sqrt{\gamma_j}\zeta_j(\mathbf{x}, t)$$

Here the linear unitary evolution of the wave-function, is described by:

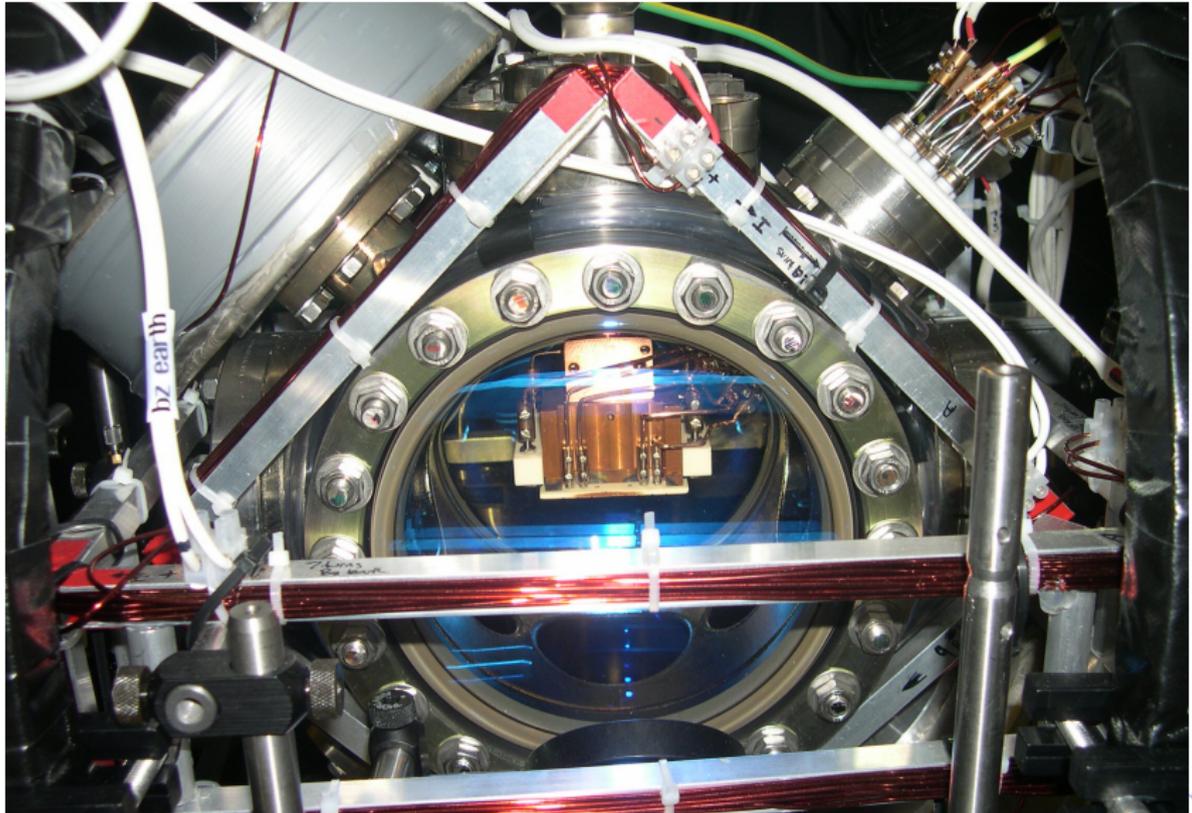
$$K_j = \hbar\nabla^2/2m - V_j(\mathbf{r})$$

while $\zeta_j(\mathbf{x}, t)$ is a complex, stochastic delta-correlated Gaussian noise with

$$\langle \zeta_i(\mathbf{x}, t)\zeta_j^*(\mathbf{x}', t') \rangle = \delta_{ij}\delta^3(\mathbf{x} - \mathbf{x}')\delta(t - t').$$

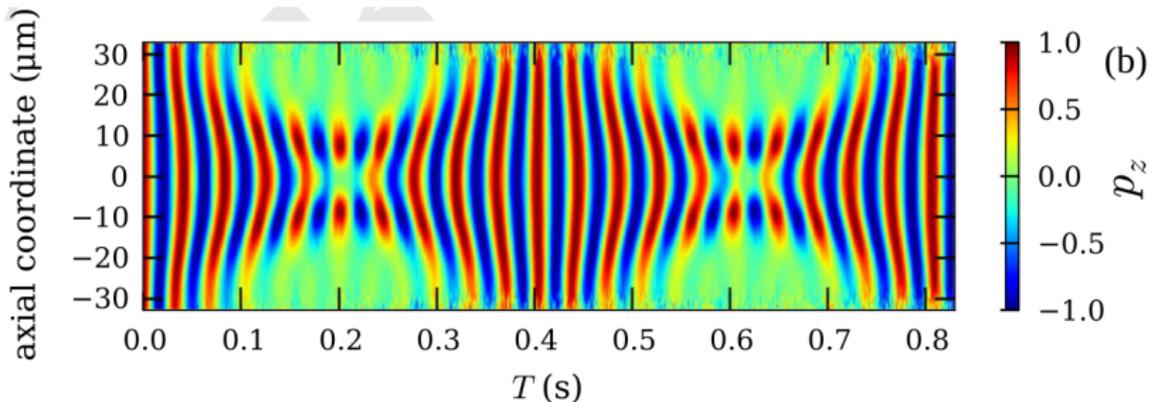
Initial fluctuations: $\langle \Delta\Psi_s(\mathbf{x})\Delta\Psi_u^*(\mathbf{x}') \rangle = \frac{1}{2}\delta_{su}\delta^3(\mathbf{x} - \mathbf{x}')$

Interferometry on an atom chip



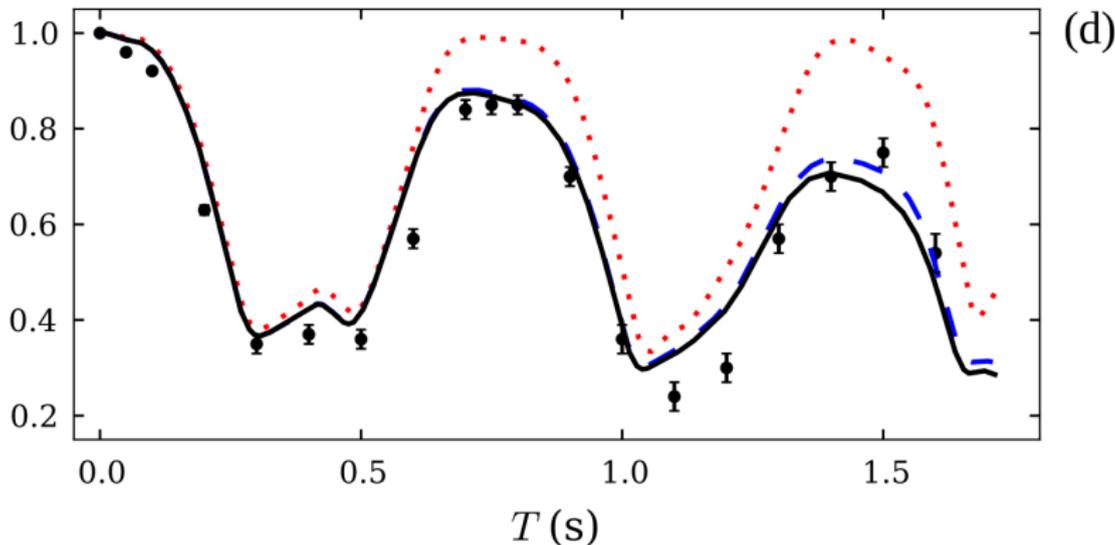
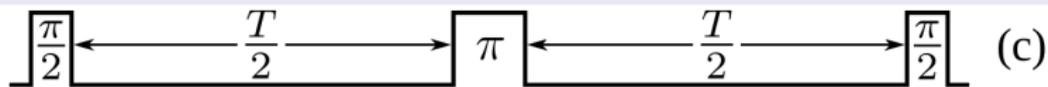
Interferometry

A two-component ^{87}Rb BEC is in a harmonic trap with internal Zeeman states $|1, -1\rangle$ and $|2, 1\rangle$, which can be coupled via an RF field.



Wigner simulations vs BEC fringe visibility

Blue line = Wigner simulation, black line = Wigner + local oscillator noise, red dots = GPE, error bars are measured



SUMMARY

Phase-space representation methods have many applications

Wigner phase-space is relatively simple!

- Maps **quantum field evolution** into a stochastic equation
- Can also be used to treat interferometry
- **Advantage:** No exponential complexity issues!
- Mathematical challenge:
 - truncation error needs to be checked: SEE Lecture 3!

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